

Markscheme

November 2018

Physics

Standard level

Paper 3

22 pages



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Section A

| Q | Question | | Answers | Notes | Total |
|----|----------|--|---|--|-------|
| 1. | а | | $m^{\frac{3}{2}}$ \checkmark | Accept other power of tens multiples of $m^{\frac{3}{2}}$, eg: $cm^{\frac{3}{2}}$. | 1 |
| 1. | b | | measured uncertainties «for one oscillation and for 20 oscillations» are the same/similar/OWTTE | | |
| | | | OR | | 2 |
| | | | % uncertainty is less for 20 oscillations than for one ✔ | | |
| | | | dividing «by 20» / finding mean reduces the random error ✓ | | |

(Question 1 continued)

| C | Question | | Answers | Notes | Total |
|----|----------|----|--|---|-------|
| 1. | С | i | Straight line touching at least 3 points drawn across the range \checkmark $ \begin{array}{cccccccccccccccccccccccccccccccccc$ | It is not required to extend the line to pass through the origin. | 1 |
| 1. | С | ii | theory predicts proportional relation $\mbox{${\tt w}$} T \propto \frac{1}{d}$, slope $= Td = \frac{c}{\sqrt{g}} = {\rm constant} \mbox{${\tt w}$} \checkmark$ the graph is $\mbox{${\tt w}$}$ line through the origin $\mbox{${\tt w}$}$ | | 2 |

(Question 1 continued)

| C | uestion | Answers | Notes | Total |
|----|---------|---|---------------------------|-------|
| 1. | d | correctly determines gradient using points where ΔT≥1.5s OR correctly selects a single data point with T≥1.5s ✓ manipulation with formula, any new and correct expression to enable g to be determined ✓ | Allow range 0.51 to 0.57. | |
| | | Calculation of g ✓ With g in range 8.6 and 10.7 «m s ⁻² » ✓ | | 4 |

| C | uestion | Answers | Notes | Total |
|----|---------|--|-------|-------|
| 2. | а | to provide a constant heating rate / power OR | | 1 |
| 2. | b | to have m proportional to $t \checkmark$ due to heat losses $\checkmark VIt$ is larger than heat into liquid» \checkmark L_{\lor} calculated will be larger \checkmark | | 2 |
| 2. | С | heat losses will be similar / the same for both experiments OR heat loss presents systematic error ✓ taking the difference cancels/eliminates the effect of these losses OR use a graph to eliminate the effect ✓ | | 2 |

Section B

Option A — Relativity

| Q | Question | | Answers | Notes | Total |
|----|----------|----|--|---|-------|
| 3. | а | | a set of rulers and clocks / set of coordinates to record the position and time of events ✓ | | 1 |
| 3. | b | i | the time in frame S' is $t' = \frac{L}{c}$ \checkmark but time is absolute in Galilean relativity so is the same in S \checkmark ALTERNATIVE 2: In frame S, light rays travel at $c + v$ \checkmark so $t = \frac{L}{(c+v)-v} = \frac{L}{c}$ \checkmark | In Alternative 1, they must refer to S' | 2 |
| 3. | b | ii | $x = x' + vt$ and $x' = L$ \checkmark «substitution to get answer» | | 1 |

| C | uesti | on | Answers | Notes | Total |
|----|-------|----|---|-------|-------|
| 4. | а | | $\frac{0.82c + 0.40c}{1 + \frac{0.82c \times 0.40c}{c^2}} \checkmark$ 0.92c \checkmark | | 2 |
| 4. | b | i | $\Delta t' = \frac{120}{0.40c} \checkmark$ $\Delta t' = 1.0 \times 10^{-6} \text{ (s)} \checkmark$ | | 2 |
| 4. | b | ii | $\gamma = \frac{1}{\sqrt{1 - 0.82^2}} = x \cdot 1.747 \checkmark$ $\Delta t = \frac{v \cdot \Delta t'}{c^2} = 1.747 \times \left(1.0 \times 10^{-6} + \frac{0.82c \times 120}{c^2}\right)$ OR $\Delta t = \frac{120}{1.747 \times (0.92 - 0.82)c} \checkmark$ $2.3 \times 10^{-6} \text{ s.s. } \checkmark$ | | 3 |

| Q | uestic | on | Answers | Notes | Total |
|----|--------|----|---|-------|-------|
| 5. | a | - | $\gamma = \frac{1}{\sqrt{1 - 0.745^2}} = 1.499 \checkmark$ $x' = \frac{1}{\sqrt{1 - 0.745^2}} = 1.499 \times (1.0 - 0) \checkmark$ $x' = 1.5 \text{ m}$ | | 2 |
| 5. | а | ii | $t' = \mathscr{C}_{vx}\left(t - \frac{vx}{c^2}\right) = x \cdot 1.499 \times \left(0 - \frac{0.745c \times 1}{c^2}\right) \mathscr{C}_{v} = -\frac{1.11}{c} \times \frac{1.11}{c}$ $\mathscr{C}_{vx}(t' = -1.1m)$ | | 1 |

(Question 5 continued)

| C | Questi | on | Answers | Notes | Total |
|----|--------|----|--|----------------------|-------|
| 5. | b | i | line through event E parallel to ct' axis meeting x' axis and labelled P ✓ | ct S' frame S frame | 1 |

(Question 5 continued)

| C | uestic | on | Answers | Notes | Total |
|----|--------|----|--|-------------------------|-------|
| 5. | b | ii | point on x' axis about $\frac{2}{3}$ of the way to P labelled Q \checkmark | Ct S' frame Q S frame | 1 |

(Question 5 continued)

| C | uesti | on | Answers | Notes | Total |
|----|-------|----|--|----------------------------|-------|
| 5. | С | i | ends of rod must be recorded at the same time in frame S' ✓ any vertical line from E crossing x', no label required ✓ right-hand end of rod intersects at R «whose co-ordinate is less than 1.0 m» ✓ | S' frame S frame S frame | 3 |
| 5. | С | ii | 0.7 m ✓ | | 1 |

Option B — Engineering physics

| Q | uestic | on | Answers | Notes | Total |
|----|--------|----|--|-------|-------|
| 6. | а | | taking torques about the pivot $R \times 4.00 = 36.0 \times 2.5$ \(| | 2 |
| | | | R = 22.5 «N» ✓ | | _ |
| 6. | b | i | $36.0 \times 2.50 = 30.6 \times \alpha $ | | |
| | | | $\alpha = 2.94 \text{ «rad s}^{-2} \text{ » } \checkmark$ | | 2 |
| 6. | b | ii | the equation can be applied only when the angular acceleration is constant ✓ | | |
| | | | any reasonable argument that explains torque is not constant, giving non constant acceleration ✓ | | 2 |
| 6. | С | i | «from conservation of energy» Change in GPE = Change in rotational KE ✓ | | |
| | | | $W\frac{L}{2} = \frac{1}{2}I\omega^2 \checkmark$ | | |
| | | | $\omega = \sqrt{\frac{36.0 \times 5.00}{30.6}} \ \checkmark$ | | 3 |
| | | | $\omega = 2.4254 \text{ rad s}^{-1}$ | | |
| 6. | С | ii | $L = 30.6 \times 2.43 = 74.4 \text{ «Js.»} \checkmark$ | | 1 |

| C | uesti | on | Answers | Notes | Total |
|----|-------|----|--|-------|-------|
| 7. | а | i | ALTERNATIVE 1: $P_{c} = P_{B} = \frac{P_{A}V_{A}}{V_{B}} \checkmark$ $= \frac{2.8 \times 10^{6} \times 1 \times 10^{-4}}{2.8 \times 10^{-4}} = 1.00 \times 10^{6} Pa \checkmark$ ALTERNATIVE 2 $2.80 \times 10^{6} \times 1.00^{\frac{5}{3}} = P_{c} \times 1.85^{\frac{5}{3}} \checkmark$ $P_{c} = 2.80 \times 10^{6} \times \frac{1.00^{\frac{5}{3}}}{1.85^{\frac{5}{3}}} = 1.00 \times 10^{6} Pa \checkmark$ | | 2 |
| 7. | а | ii | ALTERNATIVE 1: Since $T_B = T_A$ then $T_C = \frac{V_C T_B}{V_B}$ \checkmark $= \frac{1.85 \times 385}{2.8} \ll = 254.4 \text{ K. w. } \checkmark$ ALTERNATIVE 2: $\frac{2.80 \times 1.00}{385} = \frac{1.00 \times 1.85}{T_C} \ll \text{K. w. } \checkmark$ $T_C = 385 \times \frac{1.00 \times 1.85}{2.80} \ll = 254.4 \text{ K. w. } \checkmark$ | | 2 |

(Question 7 continued)

| Question | | on | Answers | Notes | Total |
|----------|---|----|---|------------------------|-------|
| 7. | b | | work done = $\langle p \Delta V = 1.00 \times 10^6 \times (1.85 \times 10^{-4} - 2.80 \times 10^{-4}) = N - 95 \text{ «J} $ | Allow positive values. | |
| | | | change in internal energy = $\frac{3}{2}p\Delta V = -\frac{3}{2} \times 95 = \mathbf{w} - 142.5 \text{ «J» } \checkmark$ | | 3 |
| | | | Q = −95 − 142.5 ✓ | | |
| | | | «-238 J» | | |
| 7. | С | i | net work is 288 – 238 = 50 «J» ✓ | | |
| | | | efficiency = « $\frac{288 - 238}{288}$ = » 0.17 √ | | 2 |
| 7. | С | ii | along B→C ✓ | | 1 |

Option C — Imaging

| C | Question | | Answers | Notes | Total |
|----|----------|----|---|---|-------|
| 8. | а | | each incident ray shown splitting into two ✓ each pair symmetrically intersecting each other on principal axis ✓ for red, intersection further to the right ✓ | For MP3, at least one of the rays must be labelled. | 3 |
| 8. | b | i | rays diverge after passing through lens OR the extension of the rays will intersect the principal axis on the side of incident rays/as if they were coming from the focal point/points in the left side/OWTTE ✓ | - | 1 |
| 8. | b | ii | by placing a diverging lens next to the converging lens OR make an achromatic doublet ✓ | | 1 |

| Q | uestic | on | Answers | Notes | Total |
|----|--------|----|--|--------------------------------|-------|
| 9. | а | | proper construction lines ✓ image at intersection of proper construction lines ✓ | objective lens objective lens | 2 |
| 9. | b | i | distance of intermediate image from objective is $\frac{1}{v} = \frac{1}{20} - \frac{1}{24}$ ie: $v = 120$ «mm» \checkmark distance of intermediate image from eyepiece is $\frac{1}{u} = \frac{1}{60} - \left(-\frac{1}{240}\right)$ ie: $u = 48$ «mm» \checkmark lens separation 168 «mm» \checkmark | | 3 |

(Question 9 continued)

| Q | uesti | on | Answers | Notes | Total |
|---|-------|----|--|---|-------|
| 9 | b | ii | ALTERNATIVE 1: | Accept positive or negative values throughout. | |
| | | | eyepiece: $m = \frac{-v}{u} = \frac{240}{48} = 5$ | The sopt positive of fregulate values throughout. | |
| | | | AND | | |
| | | | objective $m = \frac{-v}{u} = \frac{-120}{24} = -5$ | | |
| | | | Total $m = -5 \times 5 = -25$ \checkmark | | 2 |
| | | | ALTERNATIVE 2: | | |
| | | | $m = \left(\frac{240}{60} + 1\right) \times \left(-\frac{120}{24}\right) \checkmark$ | | |
| | | | <i>m</i> = −25 √ | | |

| Q | uestic | on | Answers | Notes | Total |
|-----|--------|-----|---|-------|-------|
| 10. | а | i | | | 2 |
| 10. | а | ii | to have a critical angle close to 90° ✓ so only rays parallel to the axis are transmitted ✓ to reduce waveguide/modal dispersion ✓ | | 1 max |
| 10. | b | i | long path is $\frac{12 \times 10^3}{\sin 84^\circ}$ \checkmark = 12066 «m» \checkmark «so 66 m longer» | | 2 |
| 10. | b | ii | speed of light in core is $\frac{3.0 \times 10^8}{1.52} = 1.97 \times 10^8 \text{ wm s}^{-1} \text{ wm s}^{-1}$ where $\frac{66}{1.97 \times 10^8} = 3.35 \times 10^{-7} \text{ wm s}^{-1}$ | | 2 |
| 10. | b | iii | no, period of signal is 1×10 ⁻⁸ «s» which is smaller than the time delay/OWTTE √ | | 1 |

Option D — Astrophysics

| Question | | on | Answers | Notes | Total |
|----------|---|----|---|-------|-------|
| 11. | а | | In cluster, stars are gravitationally bound <i>OR</i> constellation not ✓ | | |
| | | | In cluster, stars are the same/similar age <i>OR</i> in constellation not ✓ | | |
| | | | Stars in cluster are close in space/the same distance | | |
| | | | OR | | 2 max |
| | | | in constellation not ✓ | | |
| | | | Cluster stars appear closer in night sky than constellation ✓ | | |
| | | | Clusters originate from same gas cloud <i>OR</i> constellation does not ✓ | | |
| 11. | b | i | d=275 «pc» ✓ | | 1 |
| 11. | b | ii | because of the difficulty of measuring very small angles ✓ | | 1 |

| Q | uesti | on | Answers | Notes | Total |
|-----|-------|-----|---|-------|-------|
| 12. | а | i | $\lambda = \frac{2.9 \times 10^{-3}}{4600} = 800 \text{ mm}$ | | 1 |
| 12. | а | ii | black body curve shape ✓ peaked at a value from range 600 to 660 nm ✓ | | 2 |
| 12. | а | iii | $\frac{L}{L_{\odot}} = \left(\frac{0.73R_{\odot}}{R_{\odot}}\right)^{2} \times \left(\frac{4600}{5800}\right)^{4} \checkmark$ $L = 0.211L_{\odot} \checkmark$ | | 2 |
| 12. | b | | $M = \ll 0.21^{\frac{1}{3.5}} M_{\odot} = \gg 0.640 M_{\odot} \checkmark$ | | 1 |
| 12. | С | | Obtain «line» spectrum of star ✓ Compare to «laboratory» spectra of elements ✓ | | 2 |
| 12. | d | | red giant ✓ planetary nebula ✓ white dwarf ✓ | | 3 |

| Q | Question | | Answers | Notes | Total |
|-----|----------|--|---|-------|-------|
| 13. | а | | measured redshift «z» of star \checkmark use of Doppler formula \mathbf{OR} z~v/c \mathbf{OR} $v = \frac{\mathbf{C}\Delta\lambda}{\lambda}$ to find $v \checkmark$ | | 2 |
| 13. | b | | use of gradient or any point on the line to obtain any expression for either $H = \frac{V}{d}$ or $t = \frac{d}{V}$ or correct conversion of d to m and v to m/s \checkmark = 4.6×10^{17} «s» \checkmark | | 3 |